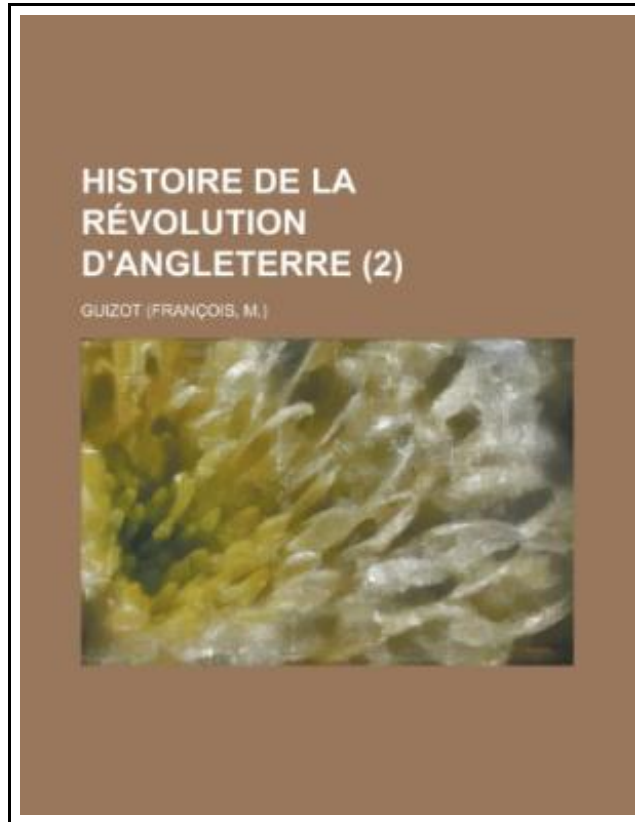


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RareBooksClub. Paperback. Book Condition: New. This item is printed on demand. Paperback. 44 pages. Original publisher: Moffett Field, Calif. : National Aeronautics and Space Administration, Ames Research Center, 1985 OCLC Number: (OCoLC)643084872 Excerpt: . . . 14 Our CO measurements add further to our understanding of this region. Figure 8 shows the MWO maps of 13CO (2-1) in the two strongest components, which correspond approximately to the dominant radio recombination line velocities. Shaver et. al. (1982) and Leisawitz and Bash (1984) find evidence in several HII regions for velocities which are anomalous with respect to their galactic position, though the differences are generally an order of magnitude smaller than the On the other hand, the difference between velocity difference between G25. 4NW and G25. 4SE. I 1 the ionized gas velocity of 59 km-s-and the strong CO component at 65 km-s-for G25. 4SE IIII is not unusual. Israel (1978) has shown that such velocity differences between compact regions and their associated molecular clouds are common, and seen in many less obscured IIII regions. The less obscured regions are presumably on the near sides of their molecular clouds, and the ionized gas is seen boiling off the cloud surface at thermal velocities. CO map that the velocity structure of the molecular gas is also complicated It is clear from the 1 I in thIS region, with the 95 km-s-cloud lying farther to the west than the 65 km-s-cloud. This offset is in the same direction as that seen in the H86a. Although the compact IIII regions are data. not at the centers of their associated molecular clouds, it can be seen from this map that the G25. 4 region has two distinct, high-column-density, molecular clouds at very different velocities that are superimposed. The...



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